

## → RIDDLE #7

### ESA's NEO Coordination Centre

#### Help Santa to save the world!

Yesterday ESA's Flyeye telescope detected a 200 meter sized asteroid that is on a collision course with the Earth. (This is fictitious!!) And not like in the riddle of last month, where the asteroid was approaching quite slowly and had a possible impact in 2049. No! This time the asteroid is coming on a retrograde orbit (inclination of 180 degrees) with an aphelion at 20 au and a perihelion of 1 au, with a predicted impact velocity of 71 km/s, pretty similar to the orbits of the Leonid meteoroids. And the predicted impact on Earth is on 24 December 2022. We have just 2 years to react.

What are our options? Let's look at Figure 1 where the Planetary Defenders have designed mitigation scenarios as function of warning time and size of the approaching asteroid. For a gravity tractor it is clearly too late, nuclear detonations are too risky and also not needed, but for civil defence (like evacuation) the asteroid is too big.

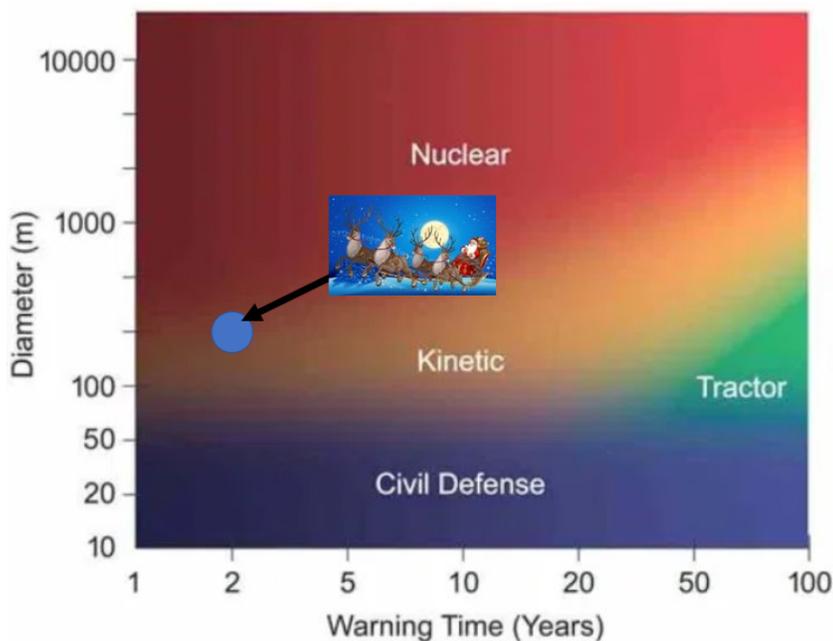


Figure 1: Planetary defence options as function of warning time and size of the asteroid. Image Courtesy of Tim Warchocki.

We need a kinetic impactor to deflect the asteroid. Unfortunately, ESA, NASA and all other space-faring nations are not yet ready to launch in such a short time a deflection mission. But Christmas time is coming up and Santa Claus has already his sledges in low-Earth orbit, ready to be loaded with the presents for all 120 million children living in Europe.

The only way to save the Earth is to load a fraction of the presents into an extra sledge and slam this sledge into the asteroid to deflect it from its fatal trajectory.

And this is the Christmas riddle: How many grams needs each of the 120 million presents be reduced if this mass is used for the required impactor? Assume a launch of the Santa Claus sledge on 24 Dec 2020 into exactly the same orbit as the asteroid, but prograde (perihelion = 1 au, aphelion = 20 au, inclination = 0 deg) and ignore the dry mass of the sledge. Assume further a circular Earth orbit with a semi-major axis of 1 au, a specific density of the asteroid of 2.0 t/m<sup>3</sup>, and assume a central impact with impulse conservation, i.e. the energy of the impactor is transferred 100 % to the asteroid. The effect of the impact is such that the new perihelion of the asteroid will be 15 000 km lower as before the impact.

## Answer

Let us first calculate the semi-major axis  $a$ , eccentricity  $e$  and orbital period  $P$  of the orbit of the asteroid and of the sledge of Santa, which are the same, just the inclination is different.

Semi-major axis	Eccentricity	Orbital period
10.5 au	0.90476	34.0 years

The impact will happen exactly after 1 year. Therefore the mean anomaly  $M$  in the orbit will be  $\frac{2\pi}{P} = 10.6^\circ$ . With an iterative algorithm we can determine the eccentric anomaly  $E = 50.7^\circ$  and with

$$\tan^2 \frac{f}{2} = \frac{1+e}{1-e} \tan^2 \frac{E}{2}$$

we calculate the true anomaly  $f = 129.5^\circ$ . And from

$$r = \frac{a(1-e^2)}{1+e \cos f}$$

we can deduce the heliocentric distance  $r$  at which the impact will happen. From the vis-viva equation

$$v^2 = \mu \left( \frac{2}{r} - \frac{1}{a} \right) \quad (1)$$

we derive the instantaneous velocity  $v$  of both asteroid and sledge, however they are in opposite directions.

We need to calculate the reduction of  $v$  such that the new perihelion is 15000 km smaller than before. Let us define  $\alpha$  as

$$\alpha = \frac{v - \Delta V}{v}$$

(e.g.  $\alpha = 0.9$  means a reduction of the velocity of 10%.)

With a reduced velocity of  $v^+ = \alpha v$  we get via the vis-viva equation the new semi-major axis  $a^+$  of the asteroid orbit after the impact:

$$a^+ = \frac{\mu r}{2\mu - v^{+2} r}$$

To get the new eccentricity, we will calculate the new semi-latus rectum  $p^+$  via the equation for the transverse component of the velocity:

$$v_t^+ = \frac{\sqrt{\mu p^+}}{r}$$



with the semi-latus rectum  $p^+ = a^+(1 - e^{+2})$ .

We know that  $v_t^+ = \alpha v_t$  and therefore

$$p^+ = \frac{(v_t^+ r)^2}{\mu}$$

This gives the new eccentricity:

$$e^+ = \sqrt{1 - \frac{p^+}{a^+}}$$

and the new perihelion radius:

$$r_p^+ = a^+(1 - e^+) \quad (2)$$

Inserting the previous equations into equation 2 yields this expression:

$$r_p^+ = \frac{\mu r}{2\mu - \alpha^2 v^2 r} \left( 1 - \sqrt{1 - \frac{(\alpha v_t r)^2}{\mu \left( \frac{\mu r}{2\mu - \alpha^2 v^2 r} \right)}} \right) \quad (3)$$

It is easy to write a small computer program to find the  $\alpha$  which will provide the requested  $r_p^+$ , however, that would be too easy. Instead, we rearrange equation 3 such that we can calculate  $\alpha$  analytically. After some manipulations equation 3 is transformed into

$$\alpha^4 \left( \frac{v^2}{\mu} (v^2 r_p^{+2} - v_t^2 r^2) \right) + \alpha^2 \left( \frac{2}{\mu r} (r r_p^+ v^2 - 2 r_p^{+2} v^2 + v_t^2 r^2) \right) + \frac{4 r_p^+}{r} \left( \frac{r_p^+}{r} - 1 \right) = 0 \quad (4)$$

The solution of this equation which we are interested in is:

$$\alpha^2 = 2 \frac{r - r_p^+}{r} \frac{\mu r_p^+}{r^2 v_t^2 - r_p^{+2} v^2}$$

Inserting the values for  $r$ ,  $r_p^+$ ,  $v$  and  $v_t$  gives  $\alpha = 0.99994$  or a  $\Delta V$  of 1.0 m/s.

Now we need to calculate the required mass of the sledge to achieve this velocity change. Conservation of momentum means ( $m_A$  is the mass of the asteroid: 8.4 million tons,  $m_S$  is the mass of Santa's sledge):

$$m_A v - m_S v = (m_A + m_S) \alpha v$$

which gives  $m_S = m_A \frac{1-\alpha}{1+\alpha}$

For the calculated  $\alpha$  the results is  $m_S = 245$  tons. Distributed over 120 million children this is just a mass of 2 gram, equivalent to less than one chocolate bonbon.

## Correct responses

This time we had one correct answer by:

- Tony Evans

Congratulations!

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